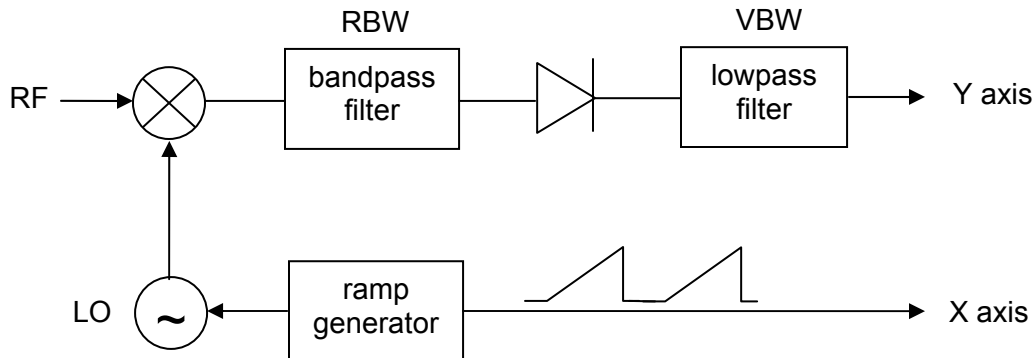
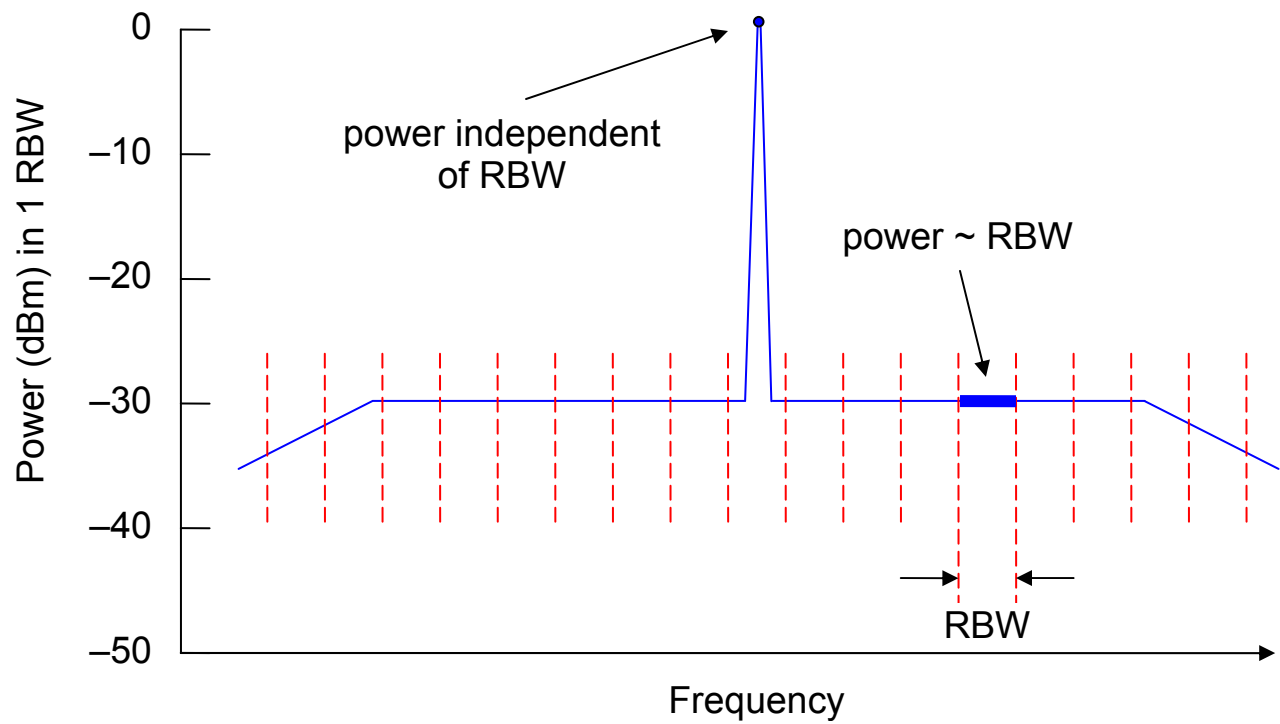


Basic parameters and functions of an analog spectrum analyzer



<i>Function</i>	<i>Description</i>
start/stop frequency	Displayed frequency range on x-axis.
center frequency/span	Ditto.
reference level	Power level at y-axis reference position, usually at top of display.
scale type	Logarithmic or linear scale on y-axis.
scale/div	Units per vertical division, usually 1, 2, 5, or 10 dB/div.
attenuation	Input attenuation. Decrease it to lower noise floor.
resolution bandwidth	Signal bandwidth of each instantaneous power measurement. Decrease it to improve sensitivity to narrowband signals.
video bandwidth	Cutoff frequency of lowpass filter following power detector. If video BW < resolution BW, output is smoothed, and sensitivity to impulsive signals is reduced.
video averaging	Display average spectrum from multiple sweeps.
sweep time	Time interval to draw trace once across screen. In order to maintain proper amplitude calibration, analyzer sets sweep time as a function of frequency span, resolution BW, and video BW.
max hold	Display max level measured at each frequency over repeated sweeps.
zero span	Configure analyzer as a fixed, tuned receiver to display power vs. time for the frequency range specified by center frequency and resolution BW.
marker	Use marker to measure frequency and power at selected points.

Resolution bandwidth



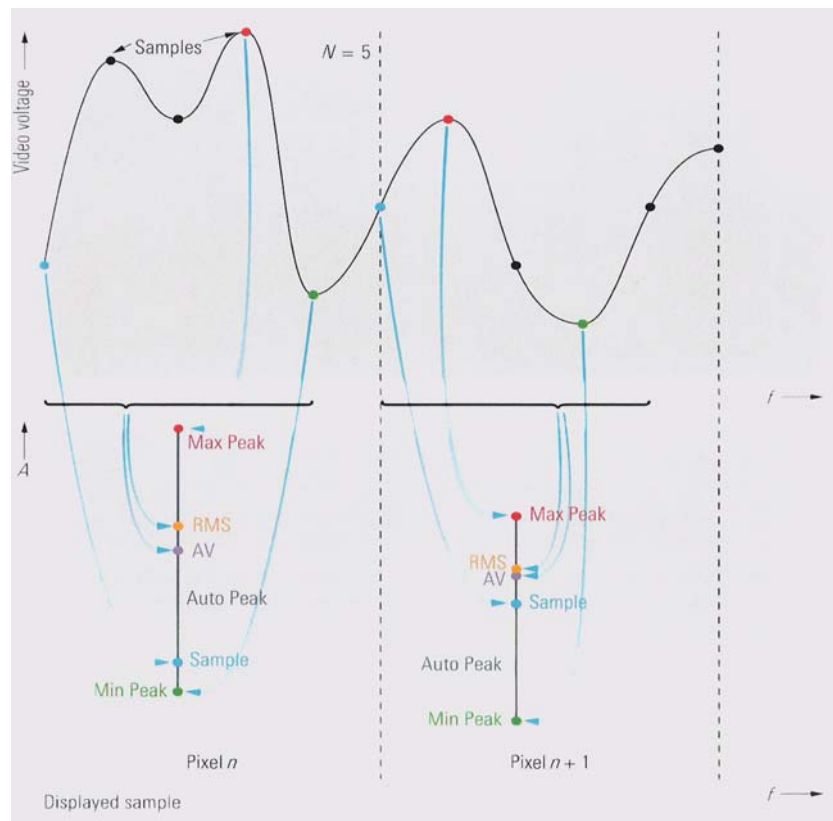
- For broadband noise with $BW > RBW$, power is proportional to RBW.
- For narrowband (e.g., CW) signals with $BW < RBW$, power is independent of RBW.

Detectors in spectrum analyzers with digital displays

Most modern spectrum analyzers use LCDs to display spectra, rather than the CRTs of older (purely analog) analyzers. With digital displays, the display resolution in both frequency and power is more limited.

The finite number of pixels in an LCD (as opposed to the effectively infinite number for a CRT) means that, for wide frequency spans and narrow RBWs, each pixel has to represent the spectral information for many sample points that lie within the frequency span of the pixel. It is left to the user to select which point to display in each pixel. Typical options include:

- *Max peak* (or simply *peak*) – the sample with the maximum power
- *Min peak* – the sample with the minimum power
- *Sample* – a “typical” value, usually either the first (in frequency) or central sample for each pixel
- *RMS* – power corresponding to the square of the RMS (root-mean-square) voltage magnitude of the samples for each pixel
- *Average* – power corresponding to the average mean voltage magnitude of the samples



From *Fundamentals of Spectrum Analysis* by Christoph Rauscher, Rohde & Schwarz, 2007

Which detector type to use depends on the nature of the signal.

- In the presence of narrowband signals, *max peak* ensures that no signal is missed.
- When only broadband, noise-like signals are present, *sample*, *RMS*, or *average* detection better represents the spectrum than does *max peak*, which will show the upper envelope of the spread in power.

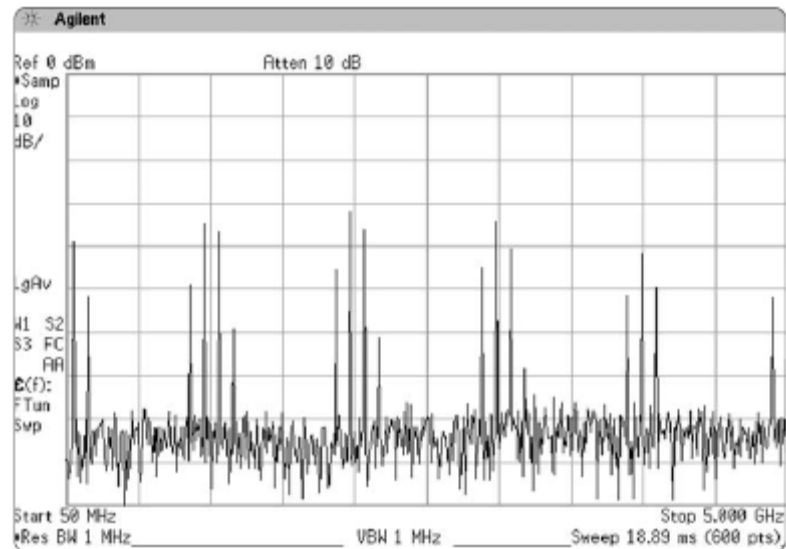


Figure 2-22a. A 5 GHz span of a 100 MHz comb in the sample display mode

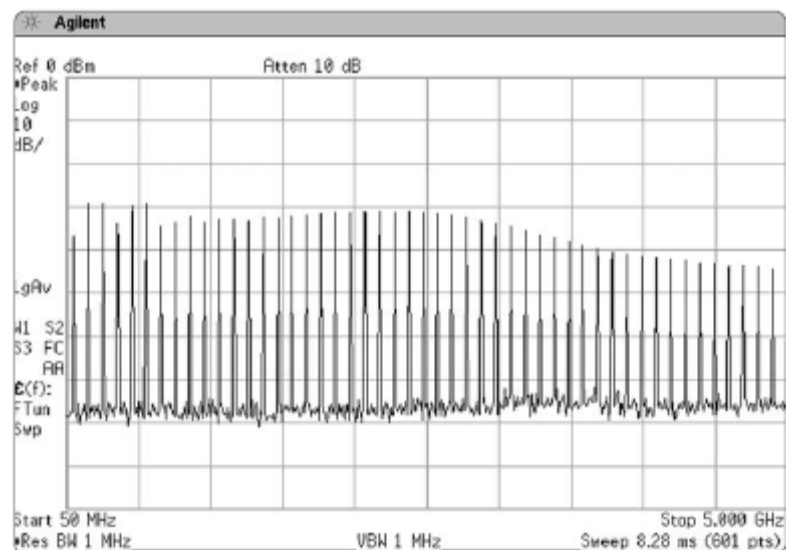
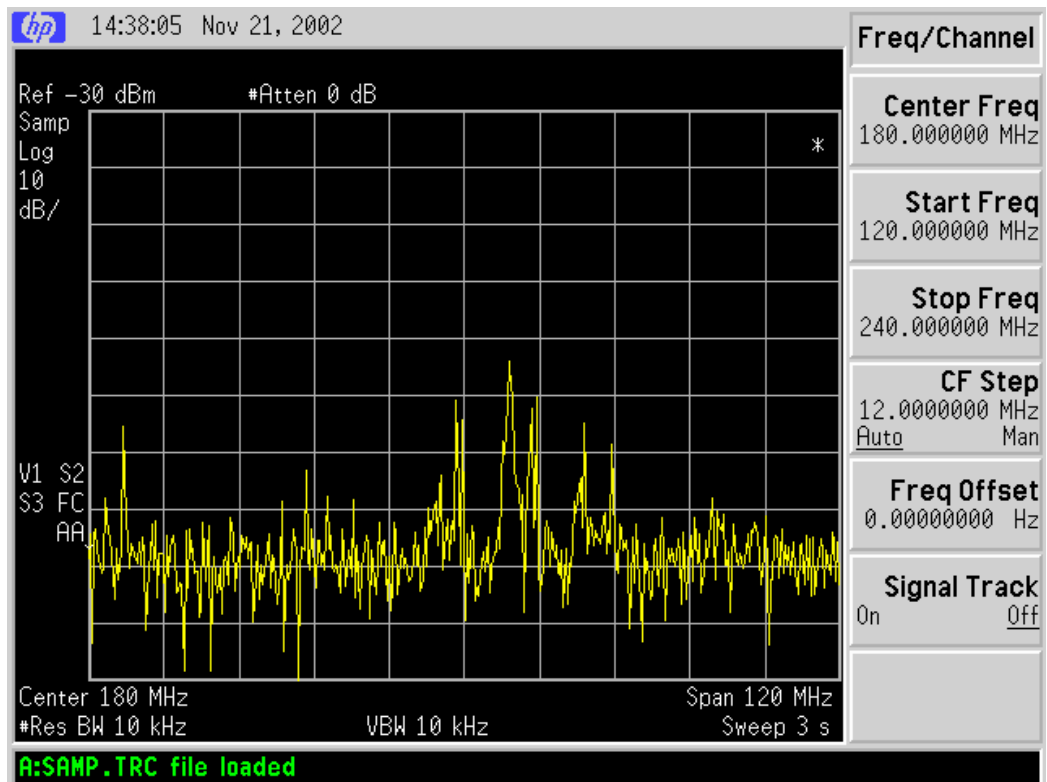
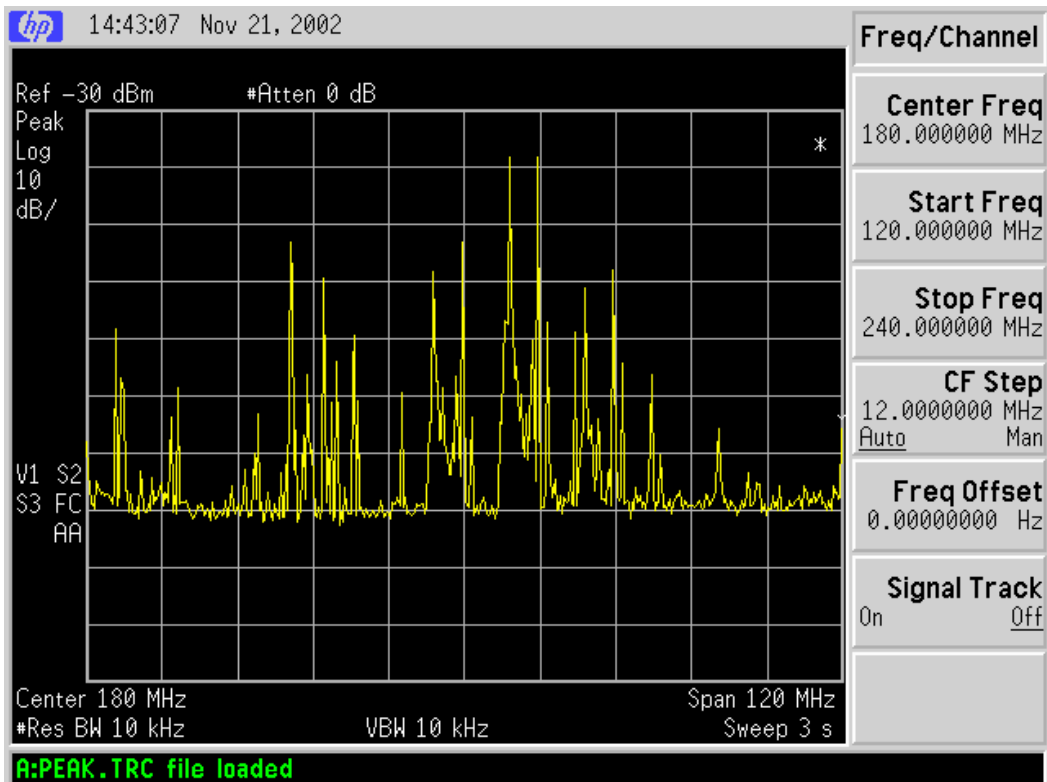


Figure 2-22b. The actual comb over a 500 MHz span using peak (positive) detection

From Agilent Application Note 150: Spectrum Analysis Basics

RFI-rich spectrum using peak (top) & sample (bottom) detection

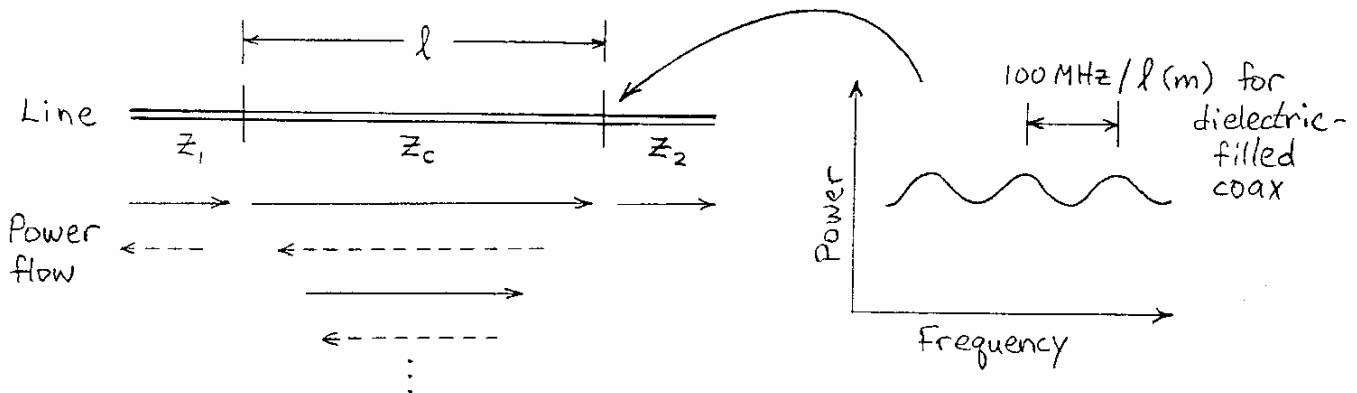


Some spectrum analyzer applications in VLBI

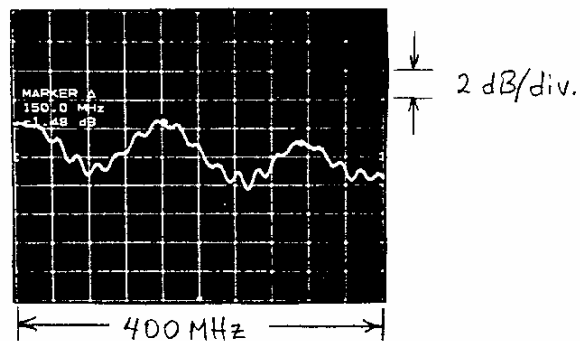
- Measure frequency response of active and passive components, e.g., filters and amplifiers.
- Measure cable loss, e.g., in cables to antenna.
- Look for ripple in broadband RF or IF signals as evidence of impedance mismatches.
- Measure LO phase noise and estimate LO phase jitter. (See “Notes on Phase Modulation of LO Signals” and IF3 LO phase noise example.)
- Test for presence of phase cal or LO modulation by measuring the carrier-to-noise power ratio of a phase cal tone and comparing against broadband power measurements. (See notes on “Using a spectrum analyzer to test for LO or phase cal modulation.”)
- Search for sideband modulation on a CW-type signal. For example:
 - 50 or 60 Hz sidebands on an LO, phase cal tone, or reference frequency signal (e.g., 5 or 500 MHz).
 - 5 kHz sidebands on phase cal tones due to MkIV cable measurement system.
 - CW sidebands on LO signal originating in phase-locked loop (e.g., 10 kHz sidebands on LO in VLBA BBC or MkIV VC).
- Search for spurious phase cal signals by turning off phase cal or unlocking the receiver LO and then looking for signals at normal phase cal frequencies.
- Search for RFI in IF or baseband signals.
- Use zero-span mode to examine temporal variability in a narrow frequency range.
- Use the analyzer as a power meter, to measure the total power over a specified frequency range, or as a frequency counter, to measure the frequency of narrowband signals.

Impedance Mismatches and Reflections

- Coax or waveguide transmission lines have constant characteristic impedance $Z_c = V/I$. ($Z_c = 50 \Omega$ is common for coax.)
- If line is terminated with active or passive device having impedance Z_c , all incident power will be absorbed without reflection.
- If device has impedance $\neq Z_c$, or if line has a break or bad connector, some power will be reflected.
 \Rightarrow Gain is decreased, and amplifier driving line may oscillate.
- If multiple abrupt impedance changes are present, multiple reflections cause ripple in power and phase spectra.



Example:
 Ripple in IF power spectrum
 due to multiple reflections over
 0.7-meter and 5-meter cable
 lengths.



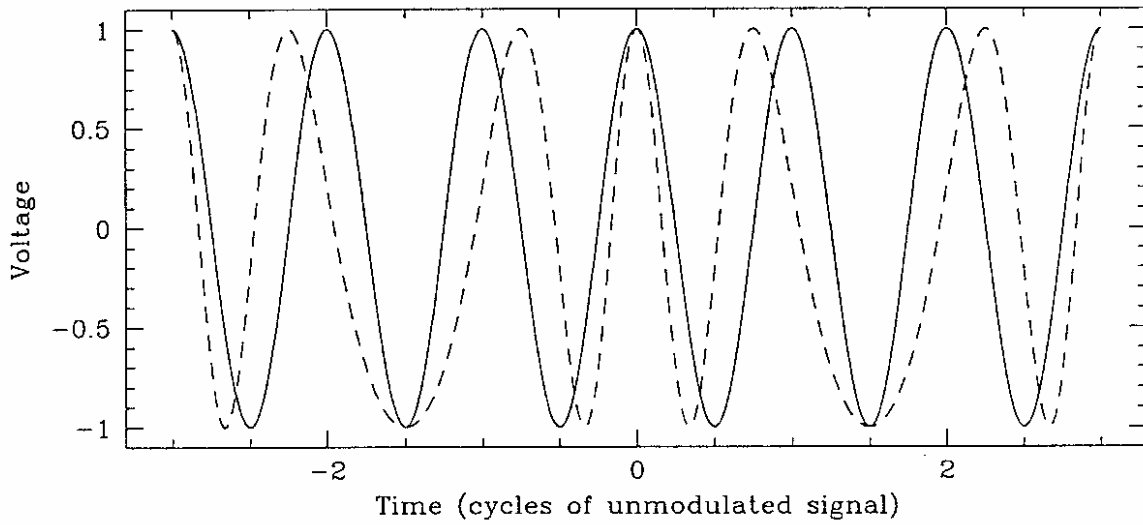
- Multiple reflections are particularly serious if they occur before the phase cal injection point, either between feed and coupler or between phase cal antenna unit and coupler.
- Even more serious are unstable, time-dependent changes in phase ripple caused by multiple reflections, which can affect measured group delay.

Phase Modulation of LO Signals

$$V_{LO}(t) = \cos[\omega_{LO}t + \phi + mod(t)]$$

Example of phase modulation in time domain:

- Unmodulated pure sine wave [$mod(t) = 0$]
- - - - Modulated sine wave with $mod(t) = (\pi/2 \text{ radians}) \times \sin \omega_{LO}t/3$



For $mod(t) = \alpha \sin \omega_m t$ and $\alpha \ll 1$ radian,

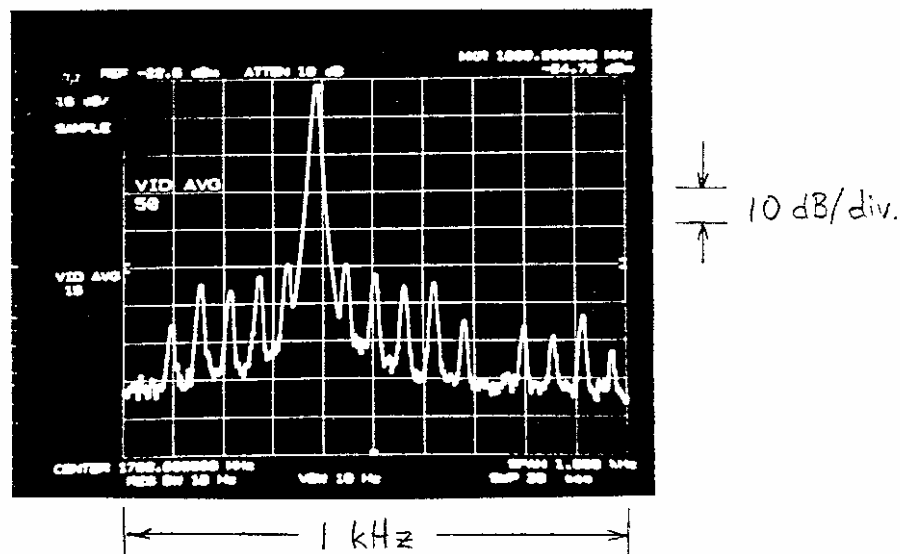
$$V_{LO}(t) \approx \cos(\omega_{LO}t + \phi) + \frac{\alpha}{2} \left[\underbrace{\cos(\omega_{LO} + \omega_m)t}_{\text{USB}} - \underbrace{\cos(\omega_{LO} - \omega_m)t}_{\text{LSB}} \right]$$

In general case, for arbitrary $mod(t)$,

- spectrum of modulated signal has upper and lower sidebands on either side of LO frequency, and
- amplitude of $\cos \omega_{LO}t$ term is reduced compared to unmodulated case.

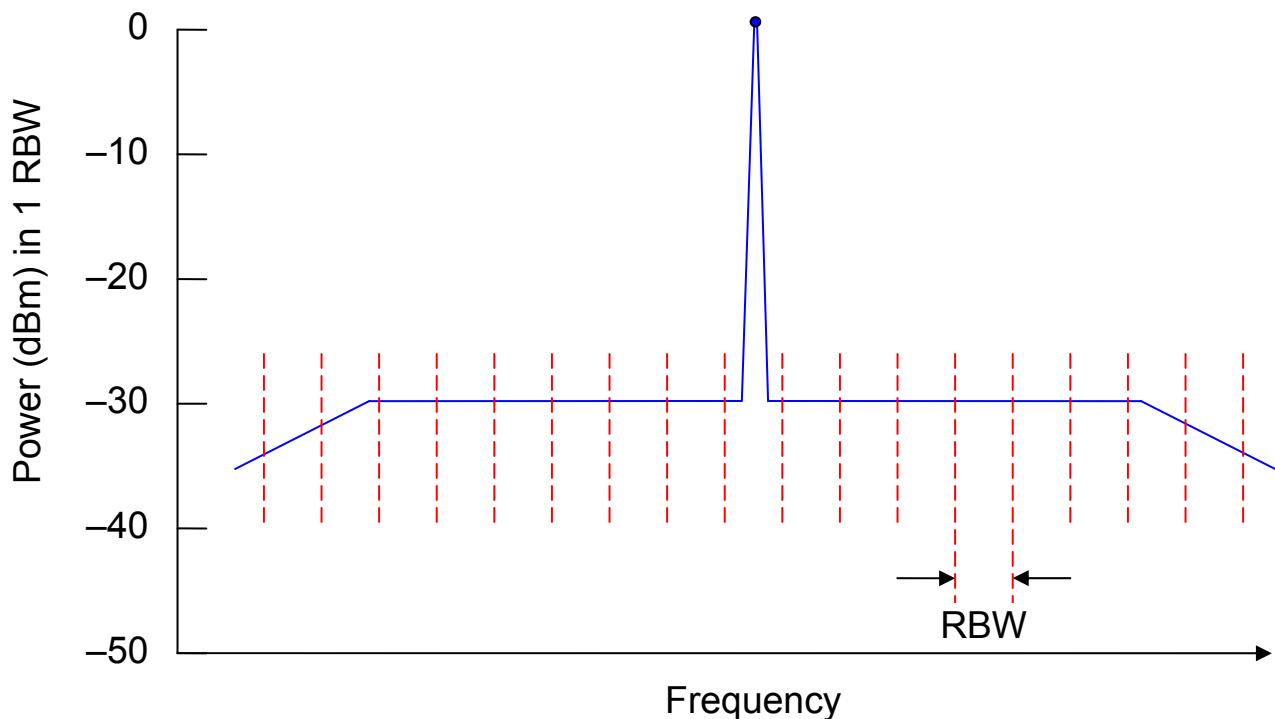
LO Phase Modulation in Geodetic VLBI

- Modulation of LO in receiver or in VC/BBC causes
 - loss of phase coherence in baseband signals relative to signals at other VLBI stations
 - degradation of VLBI sensitivity
 - shifting of signal power to modulation sidebands.
- Low-level modulation of a low-frequency LO reference signal can lead to strong modulation at high LO frequencies:
 When frequency f_1 is multiplied up to f_2 , phase noise (in degrees or radians) is multiplied by ratio f_2/f_1 , and strength of modulation sidebands is increased by $20 \log_{10}(f_2/f_1)$ dB.
 - Example: Modulation sidebands on a 5 MHz LO reference will be $20 \log_{10}(8080/5) = 64$ dB stronger at $f_{LO} = 8080$ MHz.
- A common source of modulation is 50/60 Hz hum in power supplies.
 - Example: Power spectrum of 1700 MHz LO locked to 5 MHz reference signal with weak 60 Hz modulation –



- LO modulation sidebands in VLBI systems should be > 30 dB below the LO carrier.

Measuring carrier phase noise



- Phase noise of a carrier = total power in two modulation sidebands
- RMS phase jitter of a carrier in radians = $\sqrt{(\text{power in 2 sidebands}) / (\text{power in carrier})}$

Example: Calculate carrier phase noise for spectrum above, out to the “knees” in the spectrum

Phase noise level is -30 dBm/RBW out to 6 RBWs away from carrier.

$$\rightarrow \text{power in 2 sidebands} = 2 \times (0.001 \text{ mW} / \text{RBW}) \times (6 \text{ RBW}) = 0.012 \text{ mW}$$

In practice, power is usually calculated from frequency span & RBW.

For example, if span = 12 kHz and RBW = 1 kHz,

$$\text{power in 2 sidebands} = (0.001 \text{ mW} / 1 \text{ kHz}) \times (12 \text{ kHz}) = 0.012 \text{ mW}$$

$$\text{Power in carrier} = 0 \text{ dBm} = 1 \text{ mW}$$

$$\begin{aligned} \text{RMS phase jitter of carrier} &= \sqrt{0.012 \text{ mW} / 1 \text{ mW}} \text{ radian} \\ &= 0.11 \text{ radian} = 6 \text{ degrees} \end{aligned}$$

Notes on Phase Modulation of LO Signals

Brian Corey / Haystack Observatory

11 May 1998

The output from any oscillator exhibits phase noise, *i.e.*, it is not a pure, noise-free sinusoid. The level of phase noise, or phase jitter, depends on the quality of the oscillator and of any reference signal to which the oscillator is locked, and on the oscillator control electronics. Phase modulation refers to deviation of the phase from that of a pure sinusoid. Oscillator signals may be phase-modulated either by design, as in communication systems, or unintentionally, due to intrinsic oscillator noise or poor-quality reference signals, for instance. For geodetic VLBI, the LO signals should have minimal phase modulation, in order to maximize the phase coherence of the baseband signals with those from other stations.

Phase jitter

Phase modulation appears in the time domain as jitter in the signal phase. If the signal is displayed on an oscilloscope that is triggered off a phase-stable source at the same frequency (or a subharmonic of that frequency), the zero-crossing times of a phase-modulated signal will be observed to vary about some mean value. A useful quantity to measure is the *root-mean-square (rms) phase jitter* $\Delta\phi_{rms}$ defined as

$$\Delta\phi_{rms} = \sqrt{\text{average value of (phase deviation from mean value)}^2} .$$

A rough guide that is good enough for many purposes is that

$$\Delta\phi_{rms} \approx \frac{1}{4} \times (\text{peak-to-peak phase jitter}) .$$

In the frequency domain, phase modulation appears as sidebands on either side of the carrier. There is a direct relation between the size of the sidebands and the phase jitter (strictly valid only in the case of weak modulation):

$$\Delta\phi_{rms}(\text{radians}) = \sqrt{\frac{\text{total power in both sidebands}}{\text{carrier power}}} .$$

If the value of $\Delta\phi_{rms}$ estimated in this manner with a spectrum analyzer is significantly lower than the value measured directly on an oscilloscope, then you probably missed some sidebands that lie outside the analyzer frequency range over which you searched. Note that spectrum analyzers have intrinsic phase noise themselves, so sidebands observed may originate in the analyzer and not in the signal you want to measure!

There are specifications for the maximum phase noise of most LO signals in geodetic VLBI systems. For example, $\Delta\phi_{rms}$ for the LO in a Mark III/IV video converter is specified to be $< 4^\circ$ at 100–450 MHz and $< 9^\circ$ at 450–500 MHz.

To illustrate how phase jitter is calculated from the modulation sidebands observed on a spectrum analyzer, consider a carrier with symmetric noise sidebands that extend from zero to 10 kHz on either side of the carrier. Assume the sidebands are -30 dBc down (*i.e.*, 10^{-3} as strong as the carrier) when measured with a resolution bandwidth of 1 kHz. The noise/carrier power ratio is

then $2 \times 10^{-3} \times (10 \text{ kHz}) / (1 \text{ kHz}) = 0.02$, where the factor of 2 comes from the two sidebands. $\Delta\phi_{rms}$ is then $\sqrt{.02} = 0.14$ radian, or 8 degrees.

If the modulation occurs at one or more discrete frequencies (modulation bandwidth < analyzer resolution bandwidth), rather than over a broad frequency range, then the calculation is carried out without including the (sideband BW)/(resolution BW) factor. Such discrete-frequency modulation is often observed at harmonics of the line frequency or, in the case of a VC/BBC LO, at harmonics of 10 kHz.

Tests for phase modulation of an LO in a VLBI system

1. Look for phase jitter at baseband of a test signal injected into the RF or IF part of the system. (Phase cal is generally too weak to use in this way, except in the case of gross phase modulation.)
2. Use a spectrum analyzer to look for sidebands in a baseband phase cal signal or in the LO signal itself.
3. The following simple test is sensitive to modulation sidebands >1 kHz from the carrier of a phase cal or other test signal:
 - a. Set a VC/BBC LO precisely to the IF frequency of the phase cal or test signal (*not* 10 kHz away).
 - b. In the phase cal case, set the VC/BBC bandwidth to 0.5 or 0.25 MHz. (Not all VC's have the appropriate filters.)
 - c. Turn the phase cal or test signal off. If the baseband power level drops, there is measurable power in the sidebands. (A VC or BBC is insensitive to signals at its LO frequency, so in this case the carrier power does not contribute to the baseband power level when the signal is turned on.)

How to determine whether it is the LO or phase cal that is modulated

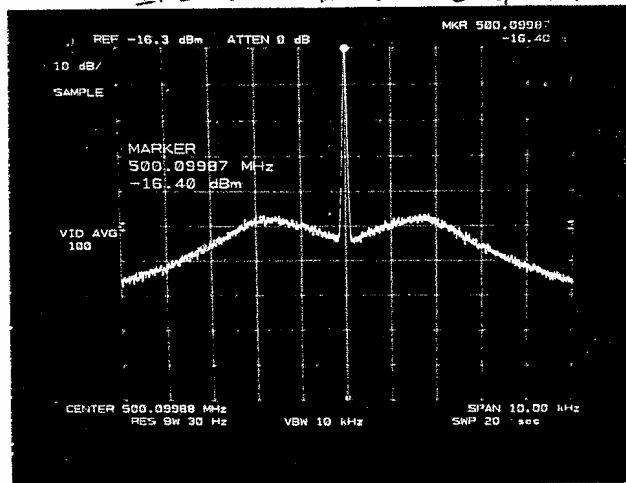
1. Inject a test signal derived from a signal source independent of the maser (*e.g.*, rubidium, cesium, or crystal oscillator), and examine the test signal at baseband. In the absence of such a source, use a separate output from the maser as the driving signal to generate the test signal. At the least, try a different 5 MHz signal from the 5 MHz distributor to drive the LO or phase cal antenna unit, and see whether the modulation changes. Note that the antenna unit will generate phase cal pulses when driven with just 5 MHz. The ground unit is needed *only* for cable measurement.
2. Look at the LO signal directly with a high-quality spectrum analyzer.
3. Try changing the LO tuning or running the LO or phase cal from a new power supply. Does the modulation change?

Beware!! The modulation may originate in the maser or its associated buffer amplifiers, in which case both the LO and phase cal signals will be modulated, and any modulation observed in a baseband or IF phase cal signal is likely to be only a small fraction of the modulation occurring

to the RF phase cal signal or receiver LO. If both the RF phase cal signal and LO signal are driven by a common modulated reference signal, the phase jitter at IF or baseband will be smaller than in the LO by the ratio $|f_{RF} - f_{LO}|/f_{LO}$ (provided the modulation frequency is lower than the loop bandwidth of any phase-locked oscillator).

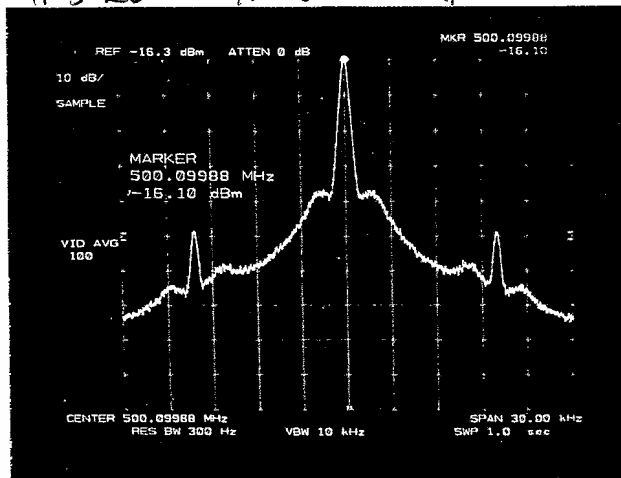
Phase noise spectra of 500.1 MHz LO in IF3 module S/N 8 -- 6 April 1993

IF3 LO - S/N 8 - 6 Apr 1993



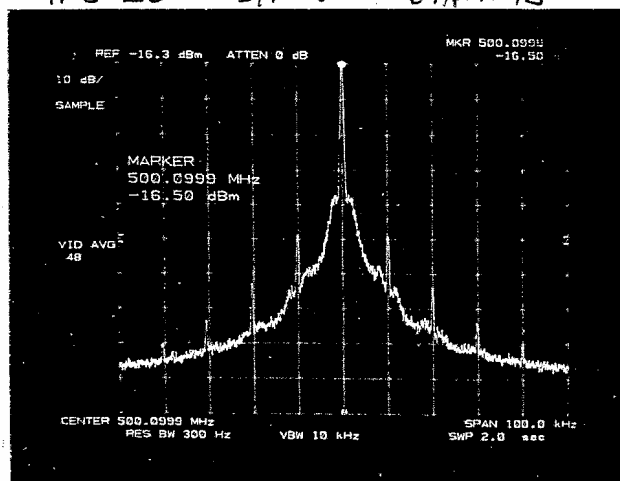
10 kHz span
30 Hz RBW

IF3 LO - S/N 8 - 6 April 93



30 kHz span
300 Hz RBW

IF3 LO - S/N 8 - 6 Apr 1993



100 kHz span
300 Hz RBW

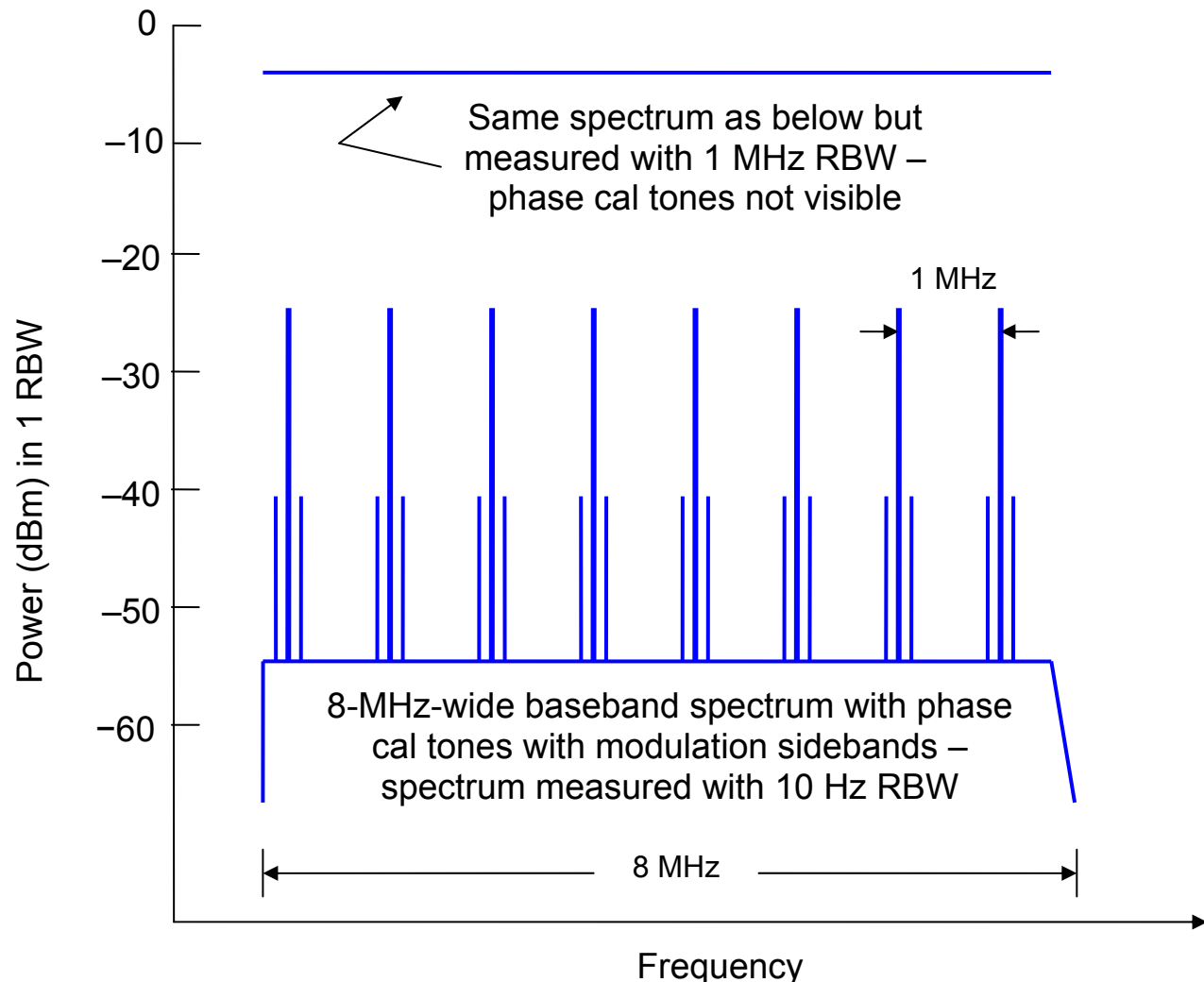
Phase noise of 500.1 MHz LO in IF3 module S/N 8
=====

6 April 1993

Freq(kHz)	dBc/kHz	dNoise	Cum.noise	rms phase (deg)
0.000	-38.8	0.000000	0.000000	0.00
1.700	-32.8	0.001124	0.001124	1.92
3.000	-40.8	0.000796	0.001919	2.51
4.000	-46.8	0.000105	0.002024	2.58
5.000	-50.8	0.000029	0.002054	2.60
6.000	-53.8	0.000013	0.002066	2.60
9.000	-54.8	0.000023	0.002089	2.62
12.000	-60.8	0.000013	0.002101	2.63
15.000	-67.8	0.000003	0.002104	2.63
20.000	-70.8	0.000001	0.002105	2.63
30.000	-75.8	0.000001	0.002107	2.63
40.000	-79.8	0.000000	0.002107	2.63

Noise spectrum measured on HP8568A spectrum analyzer.

Indirect detection of severe phase cal modulation



Measure the strength of phase cal relative to system noise in two ways:

1. Measure height of phase cal tone relative to broadband system noise with a narrow RBW. Typically phase cal is about 30 dB above noise with a 10-Hz RBW. From this measurement, calculate ratio of phase cal tone power to system power over a 1-MHz bandwidth. Value should be approximately -20 dB, i.e., phase cal is ~1% of total system power.
2. Using total power detector, measure change in baseband power level as phase cal is turned on and off.

Compare the estimates of the pcal/system power ratio from the 2 methods. If the first ratio is smaller than the second ratio by >1 dB, phase cal power is probably being lost to modulation sidebands, which might not otherwise be easily observed due to their unknown frequency range and structure.

Using a spectrum analyzer to test for LO or phase cal modulation

Modulation of a phase calibration or LO signal due to a malfunctioning power supply or other problem can sometimes be severe enough to reduce the amplitude of the signal by > 1 dB. The following technique can be used to detect such losses.

1. Display a phase cal tone at IF or baseband frequency on a spectrum analyzer, and make two measurements:
 - Measure the power level at the peak of the tone. Call it P_{pcal} .
 - Measure the power level of the noise floor surrounding the tone. Call it P_{noise} . The measurement should be done with a precision of < 1 dB. Achieving this low an error may require using either video averaging or a very narrow video bandwidth. For this to be a valid test, the noise floor must be due to the incoming signal and not to internal analyzer noise.
2. Let B_{res} be the resolution bandwidth with which the noise floor measurement was taken. Also, let r_{pcal} be the pulse repetition rate of the phase calibration system (typically 1 MHz, although the DSN, for instance, often uses other values). Using the results from step 1, calculate the fractional system power f_1 due to phase cal:

$$f_1 = \frac{P_{pcal}}{P_{noise}} \times \frac{B_{res}}{r_{pcal}}$$

On a dB scale,

$$f_1(\text{dB}) = P_{pcal}(\text{dBm}) - P_{noise}(\text{dBm}) + 10 \log B_{res}/r_{pcal} .$$

Example: For $P_{pcal} = -45$ dBm, $P_{noise} = -75$ dBm, $B_{res} = 10$ Hz, and $r_{pcal} = 1$ MHz, $f_1 = -20$ dB, or 0.01.

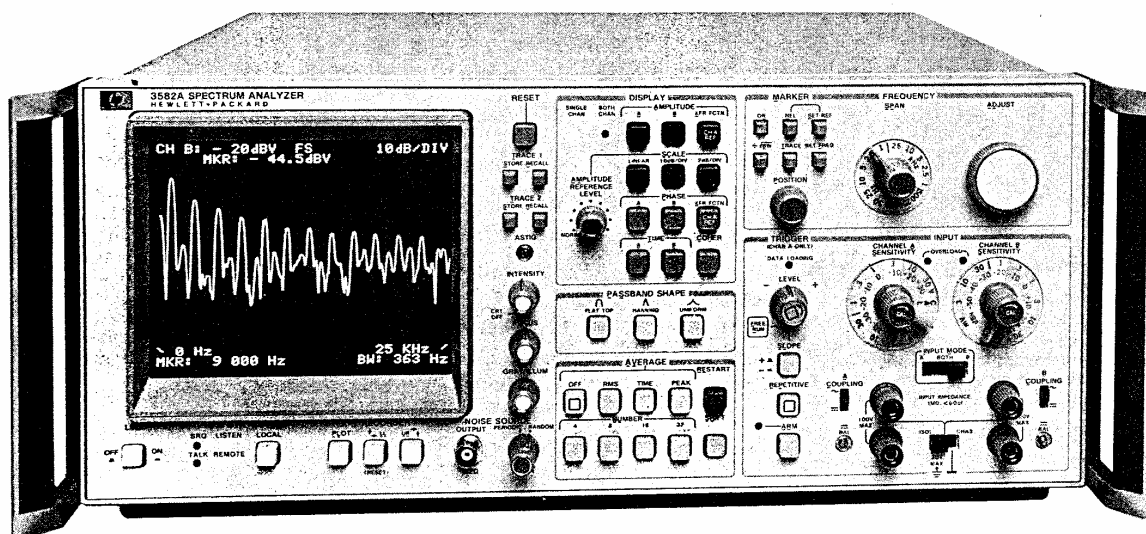
3. Using a square-law detector connected to the same IF or baseband signal (e.g., TPI detector in a VC/BBC), measure the fractional change in power f_2 as the phase cal system is turned on and off: $f_2 = P_{on}/P_{off} - 1$. f_2 is typically in the range 0.01 – 0.02. The measurement accuracy on f_2 needs to be better than $\sim 25\%$, or 1 dB, so it may be necessary to take repeated readings.
4. Compare f_1 and f_2 . Ideally, f_1 and f_2 should agree.
 - If f_1 and f_2 agree to within 25% (1 dB), then there is no severe loss of signal power.
 - If f_2 is greater than f_1 by more than 25–50% (1–2 dB), then not all of the phase cal power is going into the tones spaced r_{pcal} apart. Instead, it is likely that severe modulation of either the LO or phase cal signal has caused power in that signal to be lost to modulation sidebands. Closer examination of the IF or baseband signal may reveal the location of the sidebands.
 - If f_1 is greater than f_2 by more than 25–50% (1–2 dB), then there was probably an error in the measurements!

SIGNAL ANALYZERS

Dual-Channel, Dynamic Signal Analyzer 0.02 Hz to 25.5 kHz

Model 3582A

- Transfer function magnitude and phase measurements
- Coherence function measurement
- Phase spectrum measurement
- Transient capture and frequency domain analysis
- Internal periodic and random noise source
- Band selectable analysis for 0.02 Hz resolution
- Alphanumeric CRT annotation and marker readout



Description

The 3582A is a powerful dual-channel, real-time spectrum analyzer that solves bench or systems measurement problems in the frequency range of 0.02 Hz to 25.599 kHz. Sophisticated LSI digital filtering combined with microcomputer execution of the Fast Fourier Transform (FFT) provides exceptional measurement capability and performance.

Exceptional Frequency Resolution

The ability to resolve closely spaced spectral components is often critical in the study of subtle phenomena such as structural transfer functions. Unlike conventional dynamic signal analysis which extends from DC to some maximum frequency, the Model 3582A can "zoom in" to analyze any selected band of frequencies with dramatically improved resolution. The start or center frequency of the 5 Hz to 25 kHz band analysis spans can be adjusted in 1 Hz increments to cover the entire frequency range of the instrument. This provides resolution down to 20 milliHertz across the entire range for spectrum analysis or 40 milliHertz for transfer functions, representing as much as 5000 to 1 improvement over conventional "baseband" analysis.

Excellent Low Frequency Coverage

Many electrical and physical measurements have significant spectral information in the audio and sub-audio range. With frequency ranges from 25 kHz down to 1 Hz full scale, the Model 3582A is extremely well suited to these types of measurements. The display shown in fig. 1 represents the phase noise of a frequency synthesizer over the range of 0 to 1 Hz with a frequency resolution of 6 milliHertz.

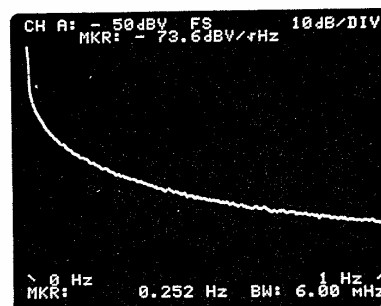


Figure 1: Phase Noise Measurement

Real Time Measurement Speed

Long measurement times can be a major limitation of swept low frequency spectrum analyzers. In high volume testing or in applications requiring substantial on-line tuning these long measurement times are both expensive and inconvenient. Since the Model 3582A uses an advanced microcomputer to execute the Fast Fourier Transform (FFT), it can perform equivalent measurements as much as one to two orders of magnitude faster than a swept analyzer.

SDR-IP software-defined radio

<http://www.rfspace.com/RFSPACE/SDR-IP.html>



SPECIFICATIONS

Frequency Range: 0.01 - 34 MHz *
Digital Down Converter: Xilinx FPGA
Synchronization : Phase and Frequency (Opt. 01-02)
PC Interface: Ethernet 100 base-T (UDP/TCP/IP)
Filters: 120+ dB 90% Alias Free BW
Decimation Rate: Variable 40-2560 (in 10 increment)
Output Sample Rate: 32 KHz to 2.0 MHz @ 24b IQ
DDC Flatness: <0.1 dB
Dynamic Range: 105+ dB
MDS Preamp Off (500Hz): -129dBm 1.8 - 30 MHz
MDS Preamp On (500Hz): -130dBm 1.8 - 30 MHz
Analog to Digital Converter: 16bit w/ Dither
Digital to Analog Converter: 14bit @ 200 MHz
Preselection: 10 Filters
Attenuators: 0, -10dB, -20dB, -30dB
Sample Rate: 80.000 MHz
Memory: 65536 x 16 bit samples (256K x 16 option)
External Radio Control: Built-in RS-232 port
Dimensions: 210 x 70 x 180 mm
Display: 16 x 2 Character LED LCD
Power: 5 Volts DC @ 1.5 Amp **
Connections: 4 x BNC (RF In, Ext Ref, Trigger, IF Out) ,
RS-232, Pulse Out, Ethernet, Power
Compliance: CE, FCC, IC

Ethernet Bandwidth Requirements:

2.00 MHz Output Sample Rate = 64Mb/s *16bit IQ
1.33 MHz Output Sample Rate = 64Mb/s
0.50 MHz Output Sample Rate = 24Mb/s
0.10 MHz Output Sample Rate = 4.8Mb/s

Expansion:

Professional Networked SDR

The high performance SDR-IP is an internet protocol (TCP/IP) software defined receiver with 1.8 MHz bandwidth. It offers the ultimate versatility for scientific research, radio astronomy, shortwave listening and amateur radio. The SDR-IP uses a high performance 80 MHz, 16 bit ADC with both dithering and randomization for best performance. The output I/Q bandwidth is configurable via software. PC communications are handled over a 100 base-T port using 24 or 16 bit I/Q words.

The SDR-IP is a complete plug and play solution. As opposed to other SDRs in the market, there is no need for additional preamplifiers, preselectors or code development. The unit is compatible with the Moetronix SpectraVue software that will have you up and running in minutes. This software includes analog mode demodulation and I/Q capture of spectrum at a 2 MS/s sample rate. For those interested in writing their own applications, RFSPACE supplies a fully documented API for communicating with the hardware. There is no need to calculate the downconverter DSP parameters. The SDR-IP just needs the center frequency, attenuator and preselector filter settings, output sample rate and mode to begin streaming data. Other solutions require complicated USB interfacing that is platform specific. The SDR-IP utilizes straight TCP/IP and ethernet for all communications. This offers the highest possible performance due to the highly optimized ethernet drivers on PC, Mac and Linux systems.

The SDR-IP packets are fully routable. This means that the receiver can be placed at remote locations. Multiple receivers can also be combined and it's data transported over a higher data rate ethernet pipe using simple switches. In this application, each SDR-IP can be set to a different IP address or port. Multiple receivers using this architecture offer the greatest performance and versatility.

Modern & New Spectrum Analyzers for the Mk-5/VLBI2010 World

Tom Clark
NASA Goddard/NVI
k3io@verizon.net

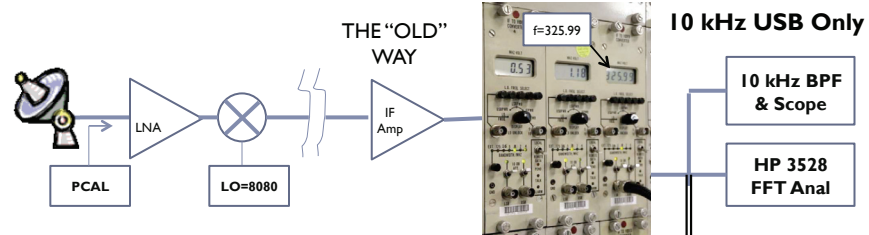
Thanks to Brian Corey & Rick Hambly for helping in my absence.

In VLBI, Spectrum Analyzers are used to:

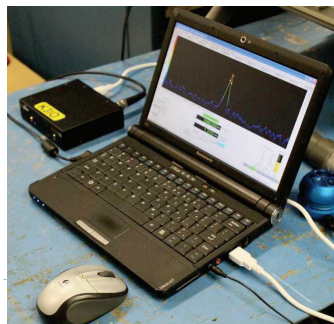
- ▶ **Detect & Identify external sources of RFI**
 - ▶ Commercial Microwave RF Analyzers (1-20 GHz)
 - ▶ often cost \$5000 - \$30,000
- ▶ **Detect & Identify local sources of RFI**
 - ▶ Most often at IF frequencies (DC-3 GHz)
 - ▶ Well matched to a new generation of SDRs (Software Defined Radios) being developed by the Intelligence Community and by Radio Amateurs
 - ▶ A major part of the SDR is implemented for cheap PCs
 - ▶ Several low cost (\$500 - \$2000) hardware platforms are available
 - ▶ Much Software comes from the Open Source, Public Domain world

This contribution will report on some of the low cost SDR options that are now available.

In the Mark-3 world we monitor Phase Cal Signals in the final baseband ("video") IF

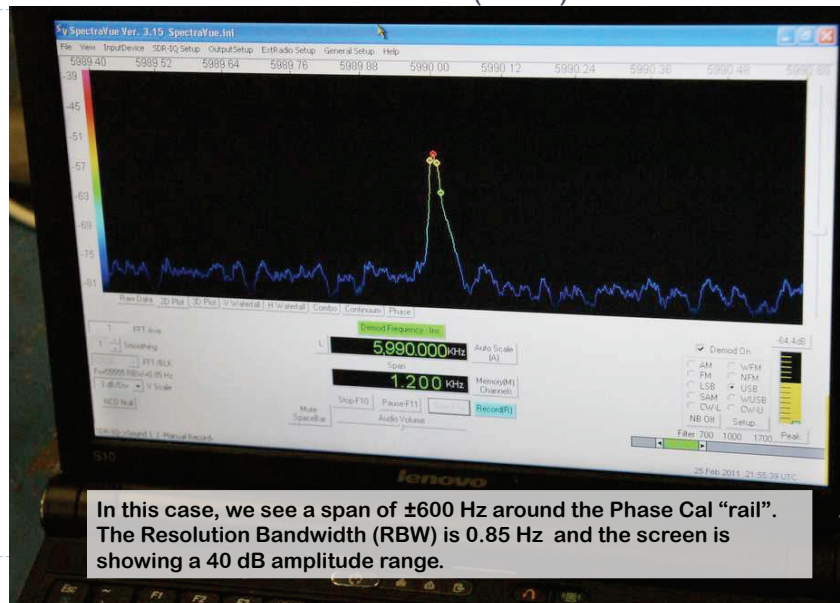


THE "NEW" WAY with \$500 SDR and a small Netbook PC

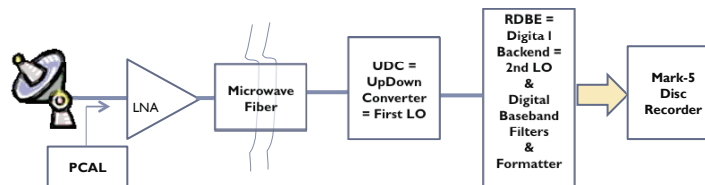


10, 1010, 2010.... kHz USB
or
990, 1990, 2990....kHz LSB
with
50+ dB Dynamic Range

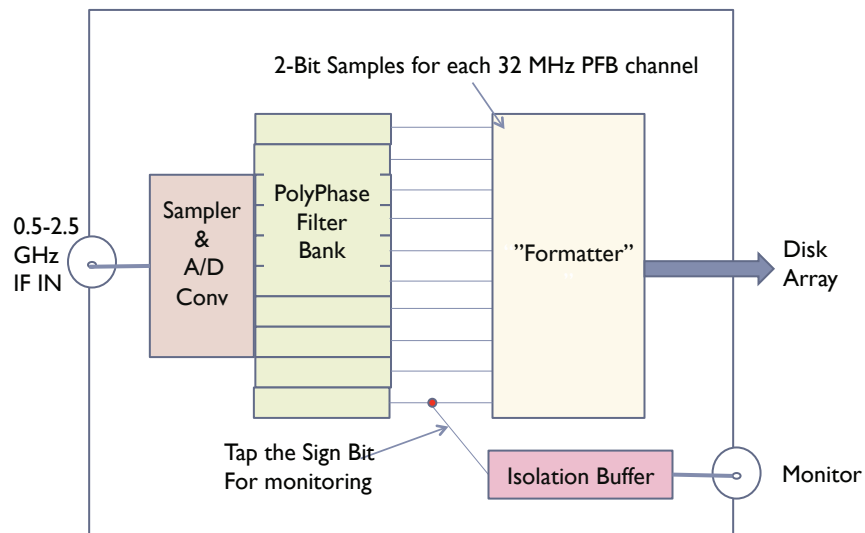
In this case, we look at the Phase Cal signal at $8080 + 325.99 - 5.990(\text{LSB}) = 8400.000$



In the new Mark-5 Digital Backend, the analog Video Converter function becomes Digital.
 The equivalent of the VC's USB/LSB BNC jacks exist only inside a Xilinx FPGA.



Real Soon Now, the RDBE Firmware will provide a digital monitor output:



Will we be able to monitor using just the Sign Bit ?

The answer is yes!

- ▶ For a “weak” signal, the S/N is degraded by a factor of $\pi/2 = 1.57 \approx 2\text{dB}$.
 - ▶ This is known as the van Vleck correction
 - ▶ The use of one-bit sampled data has been very common in Radio Astronomy
- ▶ For a strong signal, any amplitude modulation on the signal will be very distorted. An FM signal will sound perfectly normal (FM radios normally limit the signal).

Some Commercial SDRs

RF Space (<http://www.rfspace.com>) in Atlanta GA makes several SDR's:

- ▶ The \$500 SDR-IQ (used in the previous example) covers the DC-30 MHz range with up to 192 kHz bandwidth. Interface is USB. This is a competent small, cheap SDR that is a very useful piece of test equipment
- ▶ The new \$3700 SDR-IP looks like a perfect VLBI Phase Cal monitor. It's internal clocks can be locked to the Maser for fully coherent system monitoring. It interfaces via Ethernet and TCP/IP packet covering up to 2 MHz bandwidth in the 0-34 MHz frequency range. I hope to be able to try one soon.



Some (sort of) Commercial SDRs

The Funcube Dongle (<http://www.funcubedongle.com>) was developed by AMSAT-UK members to support their educational Funcube cubesat. The FCD is very interesting in that it covers the 65-1700 MHz frequency range with spans ~80 kHz wide. It is self-contained in a USB plug and costs < \$200. The FCD has (so far) been available in only limited quantities, but the ~800 users around the world have managed to make it usable on Windows, Linux and Mac computers.



Some Commercial SDRs

The \$700 “Quicksilver” QSIR (<http://qsir.wikispaces.com> or <http://www.srl-llc.com/> or <http://groups.yahoo.com/group/qsir/>) by Software Radio Laboratory in Columbus OH shows much promise. It covers DC-62 MHz (or up to ~500 MHz when oversampled) with up to 2 MHz bandwidth. The QSIR interfaces to its PC by USB. The entire design and all its support software is “open” licensed.



Some Commercial SDRs

The Italian Pegasus SDR is available for \$1000 in the US (<http://www.universal-radio.com/catalog/commrxvr/0122.html>) is quite similar to the QSIR but only covers 40 MHz.



Some Commercial SDRs

Flex Radio (<http://www.flex-radio.com/>) instrumental in introducing amateur radio to the SDR world. Most of Flex's efforts have been targeted towards full (receive + transmit) radios. Shown is their low-end Flex-1500, a \$650 DC to 54 MHz SDR that can be locked to a phase-stable external frequency standard. Flex also makes a SDR system (<http://www.flex-radio.com/Products.aspx?topic=CDRX-3200>) for the surveillance community that might be adapted to VLBI's needs.



Non-commercial SDR that may be relevant

GNU Radio (<http://gnuradio.org/redmine/wiki/gnuradio>) represents a major collaborative effort from a number of sources.

All the GNU software is supported on the USRP hardware available from Matt Ettus (<http://www.ettus.com/>). FYI – Ettus was recently acquired by National Instruments and I anticipate NI will be making a splash in the SDR world soon.



Non-commercial SDR that may be relevant

HPSDR – the High Performance SDR is an all-amateur, open- source effort (<http://openhpsdr.org/>, <http://tapr.org> and <http://iquadlabs.com/>) that has produced some **very** impressive hardware.

Rick Hambly (W2GPS) has been very active in HPSDR and can answer questions.

